



# EXOZODI Project

## ANR Programme Blanc 2010

Project coordinator : J.-C. Augereau (IPAG)

## Overview

- **Two partners:**
  - IPAG, scientific coordinator: J.-C. Augereau
  - LESIA, scientific coordinator: P. Thébault
- **Duration: 4 years** : Dec. 20, 2010 – Dec. 19, 2014
- **Budget: 430 kEuros**
  - 3 two-year post-doc positions + 6 month short term contract.  
About ~ 3/4 of the overall project budget
  - Modest contribution to near-IR interferometric instruments
  - Missions
- **Web page:**  
[http://www-laog.obs.ujf-grenoble.fr/~augereau/Site/ANR\\_EXOZODI.html](http://www-laog.obs.ujf-grenoble.fr/~augereau/Site/ANR_EXOZODI.html)



# Overview

	Budget	Permanent staff (persons.months)	Non-permanent staff (persons.months)	Overall Cost of the Project
IPAG	~ 257 kEuros	74.5	60	~1.2 Meuros
LESIA	~ 173 kEuros	72.8	56.5	~1.1 Meuros

## IPAG

J.-C. Augereau  
H. Beust  
J.-B. Lebouquin  
J. Lebreton  
B. Lazareff  
G. Zins  
2 Post-Docs

## LESIA

P. Thébaut  
V. Coudé du Foresto  
B. Mollier  
3 IRs  
2 IEs  
1 Post-Doc  
1 CCD

## External Collaborators

O. Absil (Liège)  
D. Defrère (Bonn)  
A. Brandeker (Stockholm)  
S. Charnoz (CEA/Saclay)  
M. Kama (Amsterdam)  
R. Reche (Probayes)  
E. DiFolco (CEA/Saclay)  
J. Olofsson (Heidelberg)  
J.-P. Berger (ESO)

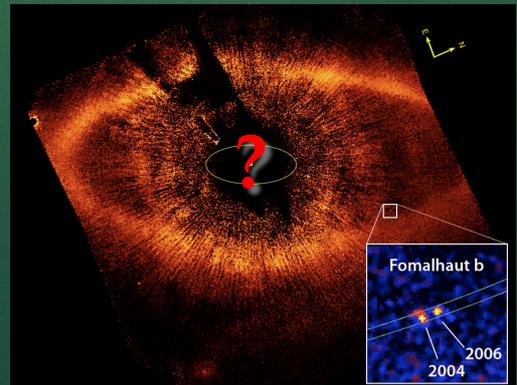
# Zodiacal dust in the Solar System

- Within ~2 AU, the **inner solar system is filled with dust** near the ecliptic plane
- Origin: **tails of comets** or when **asteroids collide**.
- It's **not a smooth cloud**:
  - Dust bands: asteroids families?
  - Dust trails: short period comets?
  - Resonant ring caused by the Earth



# Zodiacal dust in the Solar System

- **Tiny total mass :**  
equivalent to a medium-sized asteroid ( $\sim 10^{-8} M_{\text{Earth}}$ )
- But in the form of small grains,  
it is the **most luminous circumsolar component**,  
although it has a low vertical  
optical thickness ( $\sim 10^{-7}$  @ 1AU)
- Potential noise source for future  
Earth-like finding missions
- Motivated several studies  
to search for exozodiacal,  
debris disks around nearby stars



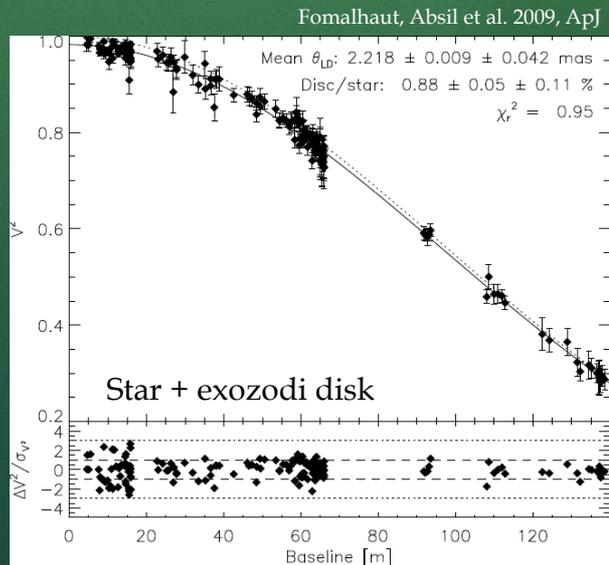
## Scientific objectives

- **Exo-zodiacal dust disks:**  
Dust clouds in and around the  
habitable zone of stars  
= warm/hot debris disks
- **Goal of the EXOZODI project:**  
detecting and understanding the  
**origin of  
exozodiacal dust**



# Scientific objectives

- Two detection techniques



# Scientific objectives

- Near-IR Interferometric Observations:
  - survey of  $\sim 100$  stars, the first of that size
  - Northern hemisphere: FLUOR/CHARA
  - Southern hemisphere: PIONIER/VLTI
- Modeling:
  - radiative transfer,
  - collisions,
  - classical N-body simulations,
  - cometary evaporation
- New numerical codes:
  - development of the next generation of model able to handle both the collisions and the dynamics
  - gas generation due to high-velocity collisions



# 5 Tasks

## COORDINATORS

- V. Coudé du Foresto [LESIA]
- J.-B. Lebouquin [LAOG]

## PARTNERS

- LAOG : PIONIER instrument team
- LESIA: FLUOR instrument team+ RI#1

## EXTERNAL COLLABORATORS

- J.-P. Berger [ESO/LAOG]

- Task 1: Instrumentation
  - Implementation of a new camera on FLUOR at the CHARA array
    - Better sensitivity and gain in statistical precision
    - 6 month short-term contract for on-site integration, testing and commissioning of the camera (at Mt Wilson)
  - Extension of the PIONIER capability to the K-band
    - Get the H-K color of exozodis, and validate some of the CHARA detections with an different instrument

# 5 Tasks

## COORDINATORS

- V. Coudé du Foresto [LESIA]
- O. Absil [Liège]

## PARTNERS

- LAOG: J.-B. Lebouquin and POST-Doc #1
- LESIA: B. Mollier

## EXTERNAL COLLABORATORS

- J.-P. Berger [ESO/LAOG]

- Task 2: observing program
  - Magnitude-limited sample:

	# MS (K < 4)	# MS w. debris (K < 4)	# MS (K < 5)	# MS w. debris (K < 5)
All	303	45	1158	103
North	156	16	536	42
South	147	29	622	61
-10° < dec < +20°	73	8	256	21

- Goal: 100 stars observed.  
~ 40 nights at FLUOR-CHARA, ~ 20 nights at PIONIER-VLTI
- Data reduction:  
Post-Doc #1, supervised by JB Lebouquin

# 5 Tasks

## COORDINATORS

• J.-C. Augereau [LAOG]

## PARTNERS

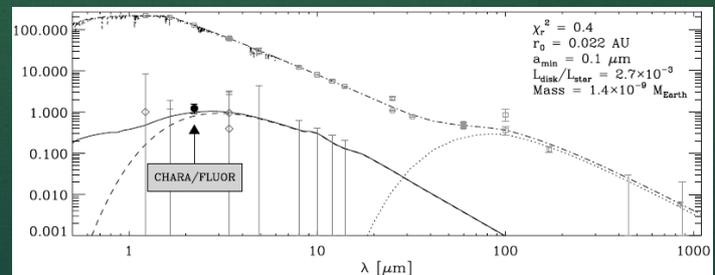
• LAOG: J.-C. Augereau, and J. Lebreton

## EXTERNAL COLLABORATORS

• M. Kama [Amsterdam]  
• J. Olofsson [MPIA]

- Task 3: SED modeling and synthetic observations
  - Making the connection between observations and dynamical models. GRaTer code
    - Improvement of the grain sublimation prescription
    - Improvement of the collision timescale calculation, and better estimates of  $dM/dt$
    - Making GRaTer compatible with dynamical models

- Expertise already available



# 5 Tasks

## COORDINATORS

• P. Thébault [LESIA]  
• H. Beust [LAOG]

## PARTNERS

• LAOG: J.-C. Augereau, and POST-DOC #2  
• LESIA: POST-DOC #3

## EXTERNAL COLLABORATORS

• A. Brandeker & R. Nilsson [Stockholm]

- Task 4: numerical simulations
  - Natural collisional evolution of asteroid belts
    - Code by P. Thébault & JC Augereau 2007
    - Update, i.p. to include grain evaporation by collisions
  - Stochastic events, such as the Late Heavy Bombardment (LHB)
    - Planet Migration
    - N-body codes
    - Comet evaporation
  - Post-Doc #2, supervised by JCA + H. Beust

# 5 Tasks

COORDINATOR  
• P. Thébault [LESIA]

PARTNERS  
• LAOG: J.-C. Augereau, H. Beust, POST-DOC #2  
• LESIA: POST-DOC #3

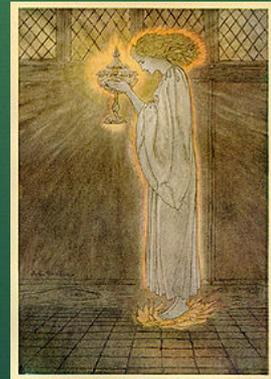
EXTERNAL COLLABORATORS  
• R. Reche [Probayes company, Grenoble]

- Task 5: 2<sup>nd</sup> generation of numerical models

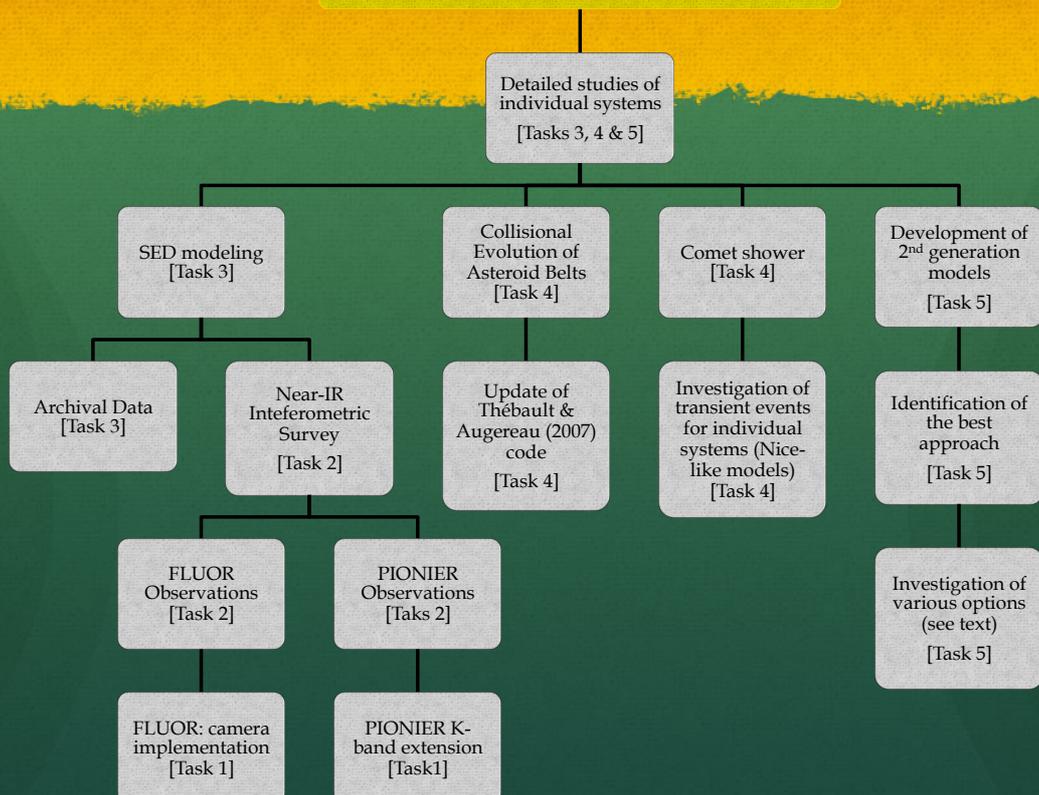
- Holy Grail

- Development of a N-body code with collisions
- Interest goes beyond exozodis: debris disks in general
- Several options to be explored

- Post-Doc #3, supervised by P. Thébault



## ORIGIN OF EXOZODIS



# Planning of the tasks

	2011				2012				2013				2014			
Task 0	Kick-Off meeting			Progress meeting #1			Progress Meeting #2				Progress Meeting #3			Closure meeting		
Task 1	FLUOR new camera implementation															
	PIONIER commissioning + K-band extension															
Task 2					FLUOR Survey											
	PIONIER pipeline				PIONIER Survey				Survey overview paper							
Task 3	Data Reduction (FLUOR and PIONIER)															
	GRaTer model improvements															
	SED fitting of individual objects with GRaTer				GRaTer compatible with dynamical codes											
Task 4	Collisional Evolution + paper writing															
	N-body simulations / transient events and comet showers															
Task 5	Exploration of methods															
	Method selection				Second-generation code development											

# Diner Tonight !



- Bouchon Montagnard,  
3 rue génissieu, Grenoble  
Tél.:04 76 27 53 50  
<http://le-bm.fr/>
- Who is coming ?