

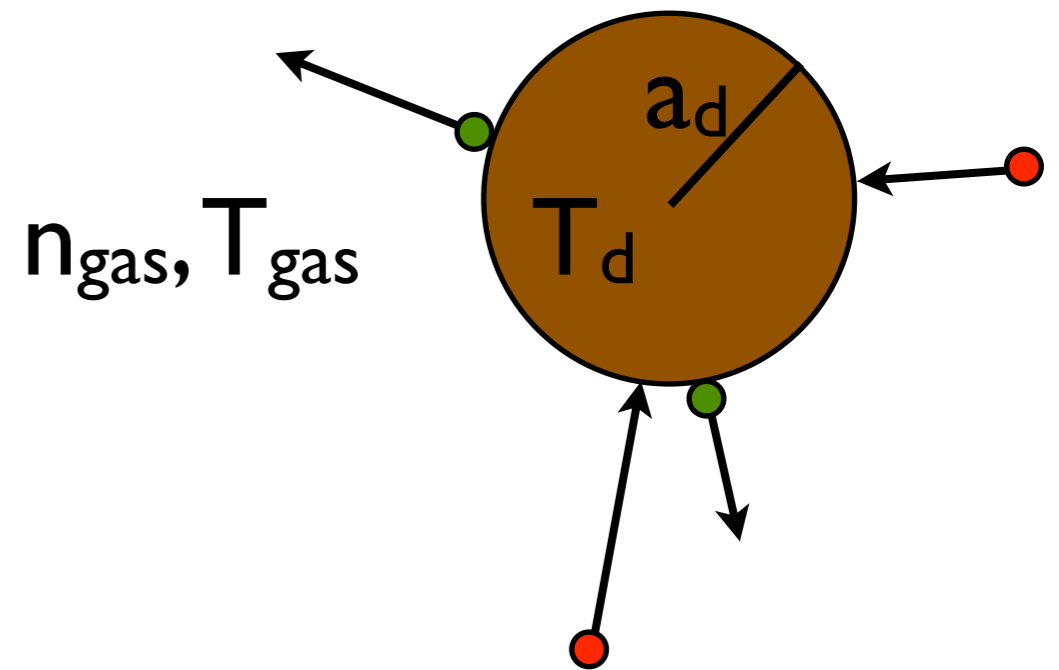
Thoughts on dust sublimation in discs

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presented by J.-C. Augereau
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The sublimation temperature



- **Evaporation** always happens.
- If $a_d = \text{const}$, **gas pressure** and evaporation balance out.
- Balance at $T_{\text{dust}} = T_{\text{subl}} = T(n_{\text{gas}}, T_{\text{gas}})$.
- For details, see Kama et al. (2009).

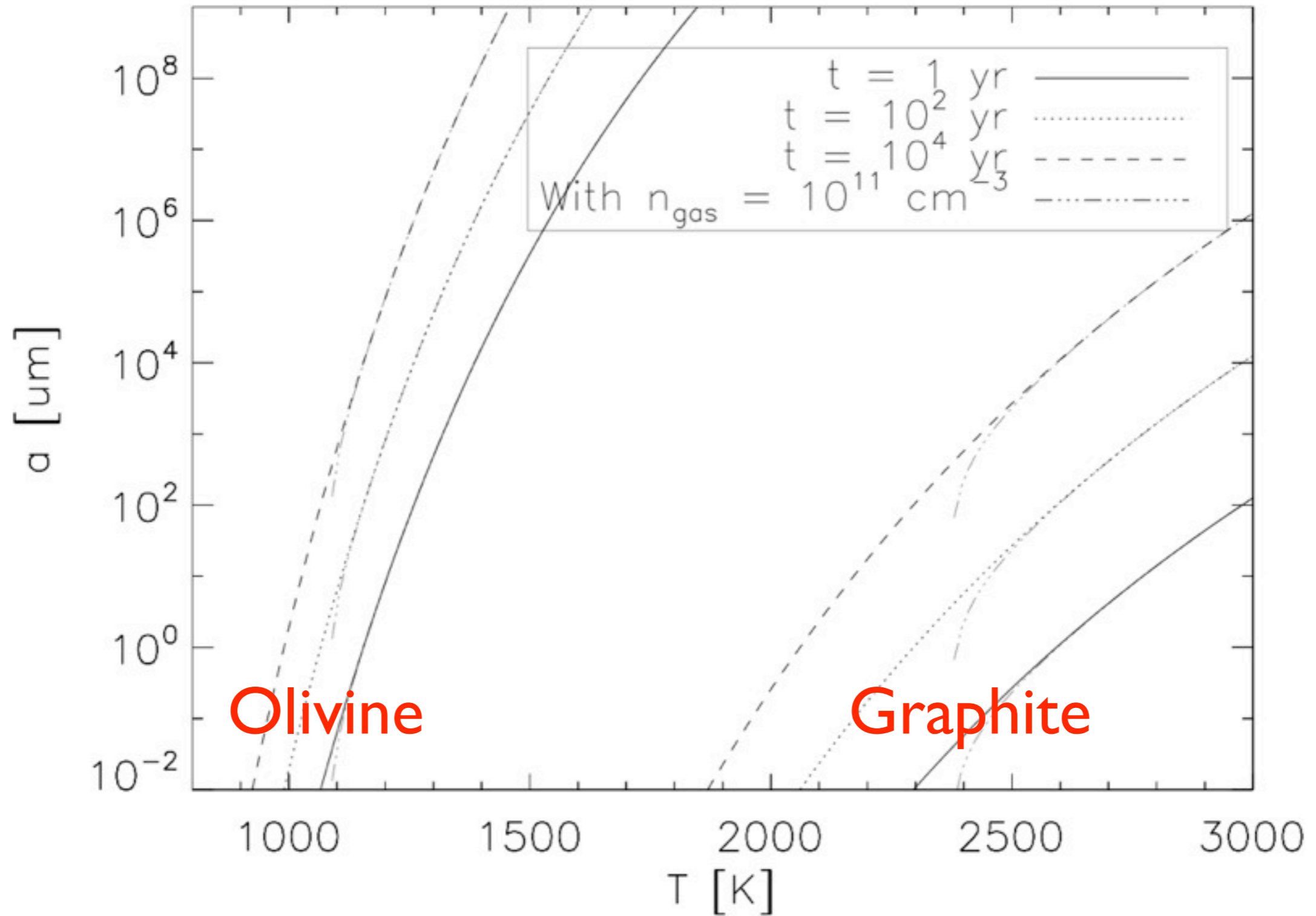
The main question

*“We assume there is **no gas** in these debris disks, and therefore the inner radial limit for the debris disk is the dust sublimation radius. For **a sublimation temperature of 1600 K** and assuming large grains in thermal equilibrium emitting as blackbodies, **the sublimation radius is...**” - Akeson et al. (2009)*

But without gas, $T_{\text{subl}} = 0$ K. So, we ask:

Which grains can survive at a given T_{dust} on timescales relevant in debris discs?

How large must a grain be to survive t [yr] at some T ?



Kama, Dominik, Augereau, in prep.

The subroutine

- For a given grain type and temperature, return the survival time.
- For a given temperature and timescale, return the surviving grains.
- Option to include gas pressure.